

Blackbody emission

As already shown, the energy density of blackbody radiation is

$$u = \frac{4}{c} \sigma T^4 .$$

I will show that the power radiated per unit area is

$$f = \sigma T^4 .$$

At equilibrium, the power radiated from a blackbody is equal to the power incident on a wall in the box containing the radiation. Thus, consider how much energy hits a wall in a time dt . The total volume available is $cdtA$, and the total energy in that volume is $ucdtA$. However, not all that energy can make it to the wall in a time dt . First of all, half of it is going away from the wall. Further, of the half that is heading toward the wall, only half will get there in time dt . Consider all the photons a distance z from the wall. To get to the wall in time dt , a photon must have $v_z dt > z$. The photons are equally likely to be going in any direction. Thus they are distributed according to the angular surface element $d\phi d\theta \sin\theta = d\phi d\cos\theta$. This says that all values of $\cos\theta$ are equally likely. But $v_z = c \cos\theta$ so all values of v_z are equally likely. As already mentioned, the half with $v_z < 0$ are immediately eliminated. When $z=0$, all the photons with positive v_z can get to the wall in time dt . When $z=cdt$, none of them can. The average over the volume out to cdt from the wall is $1/2$. Thus $1/4$ of the photons in the volume $cdtA$ get to the wall. That gives for the flux (power per unit area onto the wall)

$$f = \frac{ucdtA}{4dtA} = \sigma T^4 .$$